

Physics 324

https://javalab.org/en/category/chemistry_en/heat_en/

Particle Simulation of Thermal Conduction

https://javalab.org/en/conduction_2_en/

conduction

https://javalab.org/en/conduction_en/

convection

https://javalab.org/en/convection_en/

مخطط تغير درجة الحرارة بالنسبة للزمن عند التسخين

https://javalab.org/en/status_change_of_water_en/

heat capacity

https://javalab.org/en/heat_capacity_en/

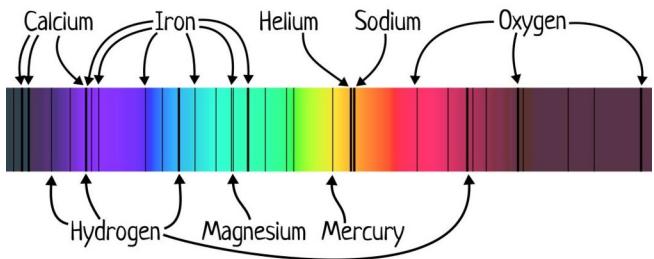
specific heat

https://javalab.org/en/specific_heat_en/

Spectral Lines of Hydrogen

[https://www.ck12.org/chemistry/spectral-lines-of-hydrogen/simulationint/Neon-Lights?
referrer=concept_details&encodedID=SCI.CHE.208](https://www.ck12.org/chemistry/spectral-lines-of-hydrogen/simulationint/Neon-Lights?referrer=concept_details&encodedID=SCI.CHE.208)

In 1814, **Joseph von Fraunhofer** studied and measured the **dark lines** in the solar spectrum. 45 years later, it was noticed that the lines coincide with the **emission lines** in the spectra of heated elements. The discovery allows us to determine the **composition of the Sun**.



Flat-Earthers are often seen saying that it is impossible to determine the **composition of the Sun** because nobody has visited the Sun before. They are wrong. **Spectroscopy** allows us to study the **composition of the Sun** and other distant celestial bodies **without going there physically**.



[https://webbtelescope.org/contents/media/images/01F8GFBXY4HW6EP5Z00C0ZZDHC?
Tag=Stars](https://webbtelescope.org/contents/media/images/01F8GFBXY4HW6EP5Z00C0ZZDHC?Tag=Stars)

[https://webbtelescope.org/contents/media/images/01F8GF9E8WXYS168WRPPK9YHEY?
Tag=Astronomy%20Basics](https://webbtelescope.org/contents/media/images/01F8GF9E8WXYS168WRPPK9YHEY?Tag=Astronomy%20Basics)

<https://webbtelescope.org/contents/media/images/01F8GF8DK2PRY4FP9DA2XPQC8S>